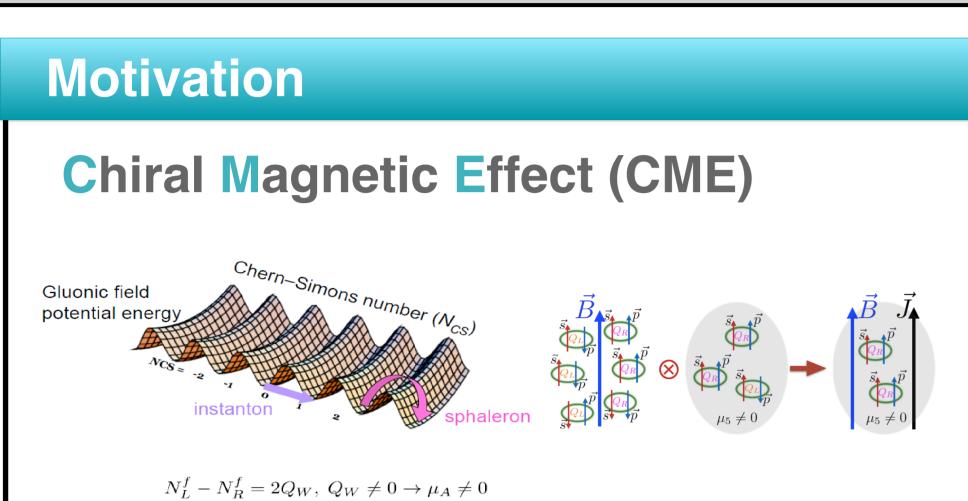


# Search for Chiral Magnetic Effect with Identified Particles in Au+Au Collisions at Jsnn = 39 GeV from RHIC/STAR

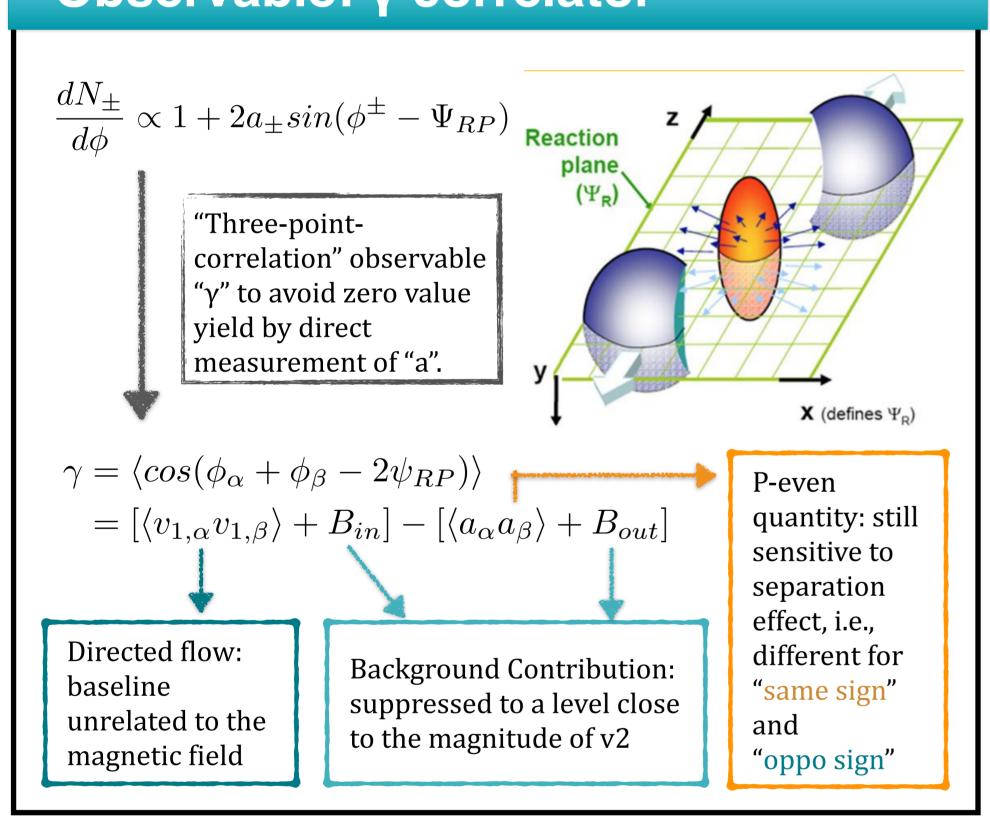
Yiwen Huang\* (UCLA) for the STAR collaboration

Chirality imbalance could occur in local domains inside the hot nuclear matter formed in high-energy heavy-ion collisions. In the presence of a strong magnetic field, this chirality imbalance will induce an electric charge separation along the magnetic field direction due to the chiral magnetic effect (CME) [1]. Previous azimuthal-angle correlation measurements [2] with unidentified charged particles have manifested charge separation signals consistent with the predictions of the CME. But the magnitudes of the background contributions have not been evaluated. In this poster, we present the correlation results with identified particles (protons, pions and kaons) using STAR data of 39 GeV Au+Au collisions. The results will be compared with those from Au+Au at  $\sqrt{s_{NN}}$  = 200 GeV, as well as the published results of unidentified particles at  $\sqrt{s_{NN}}$  = 39 GeV.



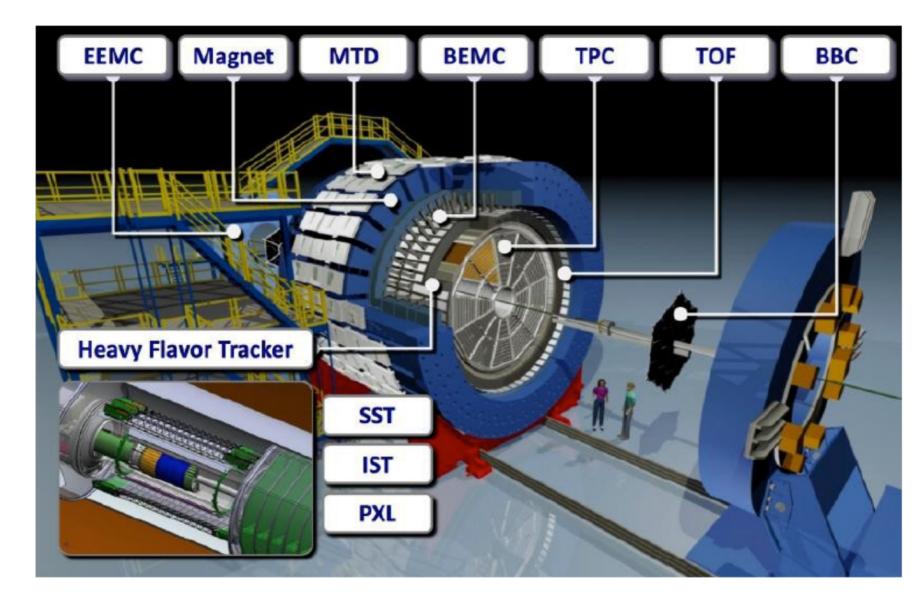
- Local chirally imbalanced domains may be created in heavy-ion collisions on an event-by-event basis. A chiral system bears a non-zero axial chemical potential,  $\mu_5$ .
- In non-central collisions, a strong magnetic field (B  $\sim 10^{17}$ - $10^{18}$  Gauss) will be produced. An electric current will be induced in chirally imbalanced domains along the B field: **Chiral Magnetic Effect** (CME).
- Charge separation fluctuations for all charged particles have been experimentally observed.
- For identified particles: Is there an indication of CME?

# Observable: y correlator



# **Experimental Approach**

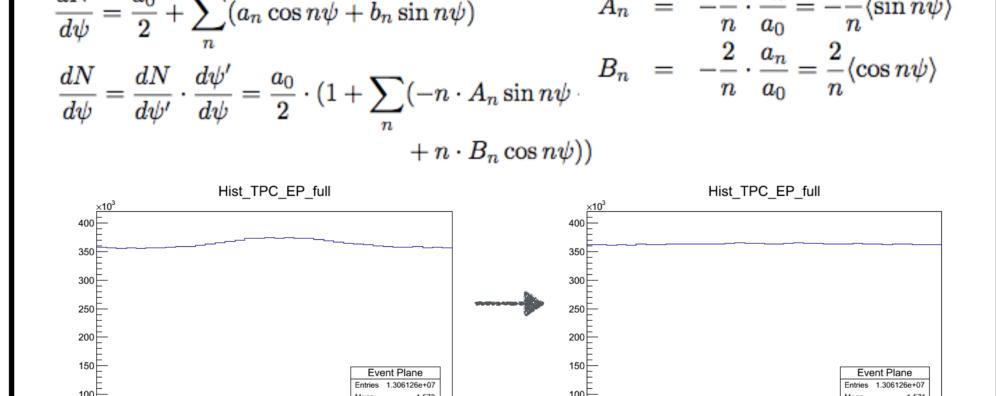
## Solenoidal Tracker At RHIC (STAR)



- Data of Au+Au at 39 GeV 200 GeV were collected by STAR detector in RHIC run 2010.
- Particle Identification:
  - Time Projection Chamber(TPC) dE/dx is used for proton/pion/kaon identification;
  - Time of Flight (TOF) detector provides mass information used for particle identification.

## **Event Plane Reconstruction**

Shifting method is used to make corrections to the event plane.



# Flow Background

- Correlation signal is contaminated with the background **contributions** due to collective motion of the collision system (flow)
  - elliptic flow coupled with transverse momentum conservation (TMC), local charge conservation (LCC) and also decay of the clusters.
- H: background-subtracted correlator, CME contributions

$$\gamma \equiv \langle \langle \cos(\phi_{\alpha} + \phi_{\beta} - 2\Psi_{\mathrm{RP}}) \rangle \rangle$$
 $\delta \equiv \langle \langle \cos(\phi_{\alpha} - \phi_{\beta}) \rangle \rangle$ 
 $\kappa_{\mathsf{K}} \equiv \Delta \gamma / (\mathsf{v}_{2}^{*} \Delta \delta)$ 
 $H^{\kappa} = (\kappa v_{2} \delta - \gamma) / (1 + \kappa v_{2}).$ 

Baseline:  $\kappa$ 

200 GeV Au+Au Ω: |η| < 1.5if higher than PHOBOS:  $(2\overline{v}_{2,\Omega} - \overline{\overline{v}}_{2,F}) / v_{2,\Omega}$ AMPT:  $(2\overline{v}_{2,\Omega} - \overline{\overline{v}}_{2,F}) / v_{2,\Omega}$ AMPT:  $\Delta \gamma / (\mathbf{v}_2 \Delta \delta)$ % Most Central

- estimated to be 1.2-1.4 [3] for three methods[4][5][6]. Indication of signal
- baseline: κ that forces  $\Delta H = 0$ κ might be different for different particle pairs: different
- possible implications 39 GeV Au+Au

% Most Central Inconclusive to claim a charge-separation signal in protonpion or in kaon-pion because signal + background is consistent with background at this level of precision.

**STAR** preliminary

## Summary

- Identified particle results for p- $\pi$  and K- $\pi$  behave as expected: Opposite sign signal is above the same sign signal, and the the signal is smaller for more central collisions.
- Cannot claim a charge-separation signal in p- $\pi$  or K- $\pi$ because the normalized signal + background value is compatible with background at this level of precision.
- Study of physics background, e.g. flow, are underway for the charged pairs.



[1]D. Kharzeev, Phys. Lett. B **633** (2006) 260. [2]L. Adamczyk et al., Phys. Rev. Lett. 113 (2014) 052302.

[3]F. Wen, L. Wen and G. Wang. arXiv:1608.03205 [nucl-th].

[4]B. Zhang, C.M. Ko, B.-A. Li and Z.-W. Lin, Phys. Rev. C 61, 067901 (2000).

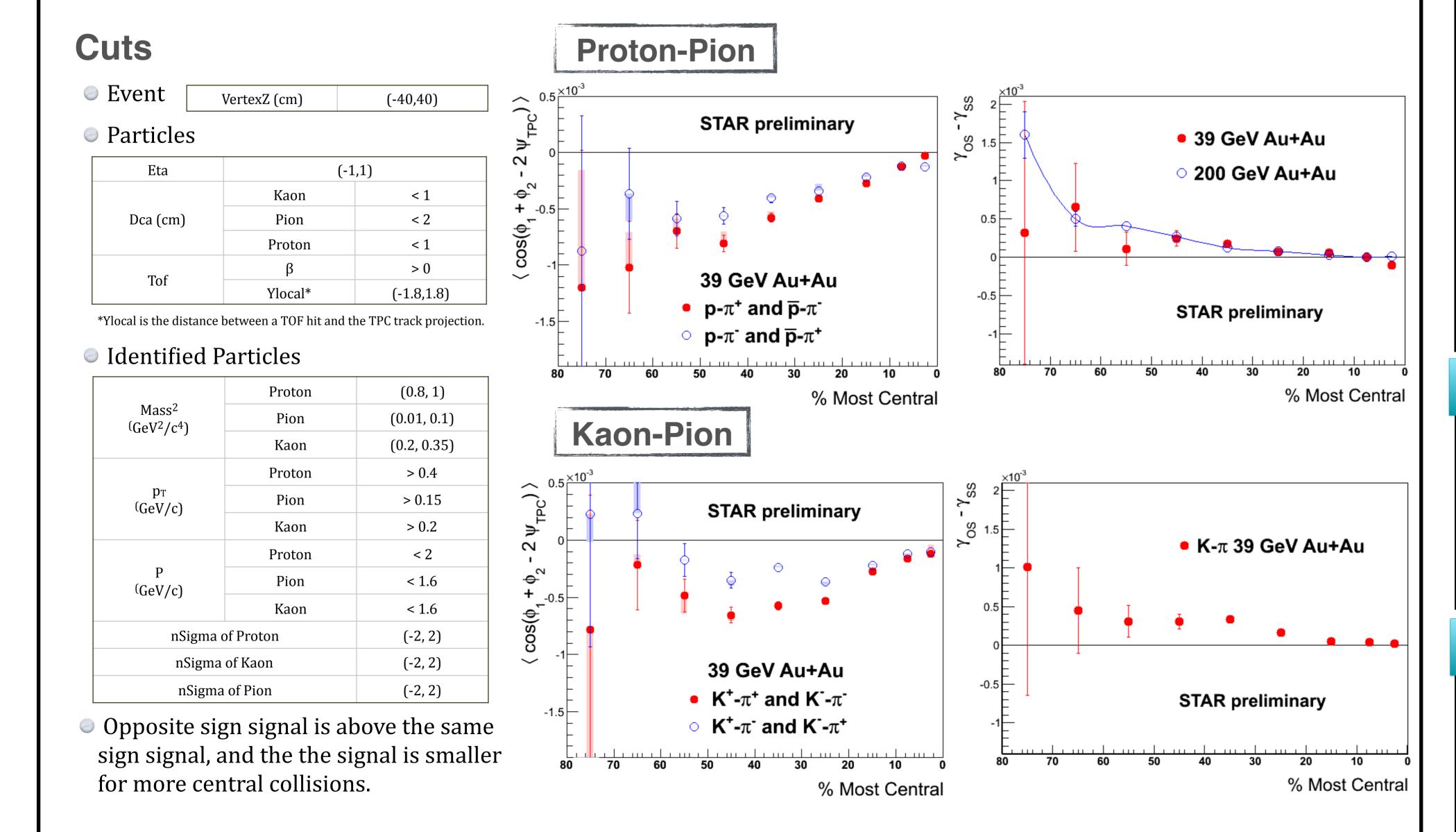
[5]B. alver et al. [PHOBOS Collaboration], Phys. Rev. C 83, 024913 (2011);

## [6]B.B Back et al. [PHOBOS Collaboration], Phys. Rev. C 72, 051901(R) (2005).

# Acknowledgements

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# **Identified Particle Analysis**



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